The consequences of tritium mix for simulated ion cyclotron emission spectra from deuterium-tritium plasmas

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What is Ion Cyclotron Emission (ICE)?

- Strongly suprathermal emission and highly spectrally structured ICRF signal spontaneously emitted from most large MCF plasmas
- Driven by energetic ion populations: fusion-born in deuterium and deuterium-tritium (DT) plasmas, also NBI and occasionally ICRH
- Spectral peaks at local cyclotron harmonics of energetic ion population
- Driven by the magnetoacoustic cyclotron instability (MCI)
 - Brought on by strong positive gradients in energetic minority particle's velocity-space



Power spectra & Frequency offsets

Using the spatiotemporal Fourier transform of the oscillatory field component, ΔB_z , we integrate over a range of wavenumbers, k, to return the power spectra as a function of frequency, Fig. 4. There are noticeable ICE peaks in all primary traces and a noticeable shifting of the dominant region $(16 < \omega / \Omega_D < 21)$ as triton concentration increases.

Using various correlation methods including the cross-correlation, phasecorrelation and shared area we defined this frequency shift by a linear relation w.r.t tritium concentration:

- distribution
- Originally detected in JET using ICRH antennae used to launch fast Alfvén wave

Why do we measure ICE?

- By understanding ICE physics, we can use ICE measurements to infer features of the energetic ion distribution
- Passive and non-intrusive
- Scaled linearly with fusion reactivity in JET
- Observed at various poloidal angles around tokamak
- ICE frequency ranges are typical of antennae already used in tokamaks

Experimental setup

- ICE is observed from most large MCF plasmas, including JET tokamak [2], Fig.2, and LHD heliotron-stellarator [3], Fig. 3
- DT fusion reaction considered: -+ --- --- --- --- (3.5MeV) + (----)
- Thermal bulk (background) ions (electrons) $T_{i,e} \sim \text{keV}$

Fig. 2 The general tokamak layout, adapted from [4], with the poloidal $(\mathbf{\theta})$ and toroidal $(\mathbf{\phi})$ magnetic fields as labelled. Poloidal field generated from applied plasma current, producing an overall helical magnetic field, spiralling around the plasma.





Fig. 3 The layout of the LHD stellarator with the magnetic coils (blue) controlling the plasma (yellow). Overall magnetic field is helical around the plasma.

 $\omega_{off} / \Omega_D = (-4.74 \pm 0.34) \xi_T + (-0.01 \pm 0.16)$ (3)

Frequency shifts are described as the preferential driving of lower energy waves for heavier, more inertial, plasmas. Best fitting simulated spectra to those in Fig. 1 is that of the 11% case, according to a τ^2 fitting method [6].

Fig. 4 Power spectra of varying triton concentrations as per legend.



Energy densities & Three-ion species gyro-resonance



Fig. 6 (above) Change in energy densities per ionic species (top to bottom) of deuterons, tritons and alpha-particles through time. Triton 10³ concentration as labelled.

Fig. 7 (right) Change in energy density per-particle ratio between the deuterons and tritons, through time, with the inverse mass ratio plotted as a dashed grey line.

<u>Higher ξ_T leads to:</u>

Fig. 5 Power spectra frequency offset as a function of tritium

- Lessening of MCI growth rate increased de-energisation and time
- Overall increased energisation of bulk ions
- Peak Δu ratio in accordance to mass ratio (gyro-resonance)

2		I	1	1	I	1%
1						
0	pro					
2						
3	- - - -					
2						11%
1						1

Simulation setup & Tritium ion concentration

- 1D3V PIC code EPOCH [5] self consistently evolves Maxwell-Lorentz system of equations
- 3.5MeV α -particle distributed as ring-beam, eq. (1), where $\perp \& \parallel$ are w.r.t magnetic field B_z component

$$f_{\alpha}(v_{\parallel}, v_{\perp}) \propto \exp\left(-\frac{(v_{\parallel} - v_0)^2}{v_r^2}\right) \exp\left(-\frac{(v_{\perp} - u_0)^2}{u_r^2}\right) \tag{1}$$

Particle volumes held equal according to the numerical weighting, eq. (2), to remove linear heating across all species, indexed σ ,

$$\frac{n_{\sigma}}{N_{s\sigma}} = const \tag{2}$$

General parameters of the JET 26148 plasma: $n_e = 10^{19} m^{-3}$; $B_0 =$ 2.1T ; B $\angle k = 89^{\circ}$; $T_{i.e} = 1 keV$; $n_{\alpha}/n_e = 2 \times 10^{-3}$

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[5] Arber T D et al. 2015 Plasma Physics and Controlled Fusion 57 113001 [6] Naylor T and Jeffries R D 2006 MNRAS 373 [7] Dendy R O et al. 2023 Phys. Rev. Lett. 130 105102

By assuming both DT species $\breve{A} | \breve{A} |_{10^{\circ}}$ reach gyro-resonance through $\overline{(1)^{10^{-1}}_{10^{-1}}}$ Larmor radii matching [7], one derives the ratio between their change in energy densities perparticle as the inverse of their mass ratios, Fig. 7.



Conclusions

- First principles simulations of ICE spectra with tritium concentrations in the range 0% to 50% show best agreement with JET 26148 ICE when concentration = 11%, which coincides with the actual experimental value
- Preferential driving of shorter and slower plasma waves ۲
- An increase in the magnitude of energy transfer from the alpha-particles • to the DT bulk through slower growing MCI
- Frequency offset following a linear relation w.r.t ξ_T , eq. (3) ۲
- Gyro-resonance between DT species, Fig. (7)